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History of EISCAT - Part 5: Operation and development of the system during the first two decades

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Abstract.

15 This paper gives an inside view of the first 20 years of operation of the Kiruna-Sodankylä-Tromsø (KST) part of EISCAT as experienced and remembered by myself. The paper is subdivided into an Introduction and 13 sections. Sections 1 to 6 describe the organization, staffing and responsibilities of the Sites, with particular emphasis on the transmitter-related work at Tromsø and the commuting of staff and equipment between the Sites. The Headquarters operation is treated in Section 7. The
20 UHF radar system is treated in Section 8. Section 9 is a review of the VHF system, including a summary of transmitter and antenna problems not available elsewhere in easily accessed media. Section 10 treats the computer system and the proprietary control languages EROS, TARLAN and CORLAN. Section 11 describes the signal processing hardware, with special emphasis on the Alker correlator, its idiosyncrasies and the gradual unlocking of its capabilities through UNIPROG, the
25 GEN-system and the G2-system, culminating in the ability to run alternating codes experiments routinely. Section 12 presents the time- and frequency keeping, a non-trivial task in the early 1980s. Finally, Section 13 discusses the UHF spectrum problem and relates how the UHF system had to be constantly upgraded in order to be able to co-exist with the emerging cellphone networks until the final closure of UHF reception at Kiruna and Sodankylä in 2012. The paper ends with some
30 personal reflections.

Introduction

35 When I started as Site Manager of the Kiruna site in June 1981, I was still a green newcomer to the geophysics community and my understanding of the scientific tasks that the Founding Fathers of EISCAT and the Associate scientists had set for themselves was almost nil. In a way, the whole time from then until I departed EISCAT in 2008 came to be a continuous learning-on-the-job process, peaking during the Svalbard radar project years and culminating in my leading the
40 EISCAT_3D feasibility study. I have always regarded my role in EISCAT as that of a *machine physicist*, a concept I learned of at CERN, but I am happy that in helping to introduce the study of meteor head echoes with the UHF system I was also able to do a bit of science with the instrument...

45 Thinking back, I am extremely grateful that I was able to join the EISCAT community right at the beginning of the operational phase and given the opportunity to help developing and steering the Association through a period that was arguably its best years. In my view, EISCAT was established at the best possible time, when there existed a brief window of opportunity where all the technical prerequisites for wideband incoherent scatter studies of the ionosphere were on hand at the same time:

50 In the early 1970s, blocks of unused and partially unallocated VHF and UHF spectrum in the frequency ranges optimal for ionospheric incoherent scatter observations and wide enough to receive the full scatter spectrum (10-15 MHz), were still available in the Nordic countries. Also, the theory of incoherent scatter was well established, several ISR installations had been working for a
55 number of years and high power UHF and VHF radar technology, a product of the 1950s nuclear craze, had been developed to the point where pulsed multi-megawatt transmitters could be had from industry. Low noise UHF receivers had become commercially available, the emerging satellite communications industry had prompted the development of standardized reflector antennas in the 32-m class, digital signal processing tools capable of doing full justice to wideband incoherent
60 scatter signals were coming on line and their performance was increasing by orders of magnitude every few years - so the technology was there, albeit large, costly and partly clumsy. Last but not least, the amount of sensitive electronic devices in ordinary households was very small or

Abstract

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1. Introduction

35 EISCAT, the European Incoherent SCATter radar system, is a multi-site *incoherent scatter radar* system, originally planned for studies of the auroral ionosphere and located in the auroral zone in northern Finland, Norway and Sweden. Thanks to its ability to provide spatially and temporally resolved measurements of plasma parameters (plasma density, ion and electron temperatures, ion mass and bulk velocities) throughout the ionosphere, from the D-layer to the topside, incoherent
40 scatter is a powerful ground-based tool for ionosphere and upper atmosphere studies.

EISCAT was conceived in the late 1960s by a group of Nordic ionospheric physicists, who managed to win URSI support for the concept at the 1969 URSI General Assembly. Scientists from France, Germany and the United Kingdom then joined the initiators, the movement eventually
45 leading to the establishment of the EISCAT Scientific Association. Important lessons from an already operating incoherent radar system, personal memories of the formation process, the financing negotiations and the subsequent design and construction of the three "mainland" radar stations have been published in previous papers of the present series (see references below).

50 In retrospect, it is clear that the EISCAT system was established at the best possible time, when there existed a brief window of opportunity where all the technical prerequisites for wideband incoherent scatter studies of the ionosphere were on hand at the same time:

In the early 1970s, blocks of unused and partially unallocated VHF and UHF spectrum in the
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nonexistent, so the risk of a radar system established in a populated environment generating interference to consumer equipment was negligible.

Over its more than forty years of active operation, the EISCAT system has generated a vast amount of groundbreaking ionospheric, magnetospheric and middle atmosphere science. Starting from the "yellow book"-list of key scientific questions, the so-called "eleven wonders of EISCAT" (EISCAT Steering Committee, 1974), the research has taken off in many different directions, some aiming at investigating pre-existing hypotheses and models of the ionosphere and thermosphere, others looking for rare and hard to detect plasma conditions and processes, and yet others following up the surprising number of observations of phenomena and processes not considered in the planning stage, such as VHF and UHF PMSE (e.g. Röttger et al., 1988), non-equilibrium plasma processes at F region altitudes (e.g. Lockwood et al., 1988), meteor head echoes and different kinds of coherent echoes.

This work has been extensively reported in a large number of peer-reviewed papers, conference proceedings and internal reports; according to the statistics kept by EISCAT Headquarters the total number of publications now exceeds 2500! It is beyond the capability of a single historical paper to give due credit to all this work; it stands solidly on its own merits.

However, the first step towards all these achievements has been the generation, recording and distribution of raw radar data, and in many cases also the analysis of the data into physical parameters. This work has been the responsibility of the EISCAT staff. It may not be very visible to the typical data user, but nevertheless all-important to the overall mission. I was part of this work from 1981 to 2008, and probably because of this long history I have been asked to share my memories of what life at EISCAT was like during the first two decades. In the following, I try to give my insider's view of the life and work at the three sites and Headquarters and how the original, "mainland", radar system was commissioned, operated, maintained and developed.

The initial planning and development of the Association before the official inauguration in 1981 has already been described by five of the "founding fathers of EISCAT" (Hultqvist 2011, Oksman 2011, Holt 2012, Bauer et al. 2013 and Haerendel 2016) from their respective national perspectives. My story picks up by where those contributions end. It is by and large based on my own notes and recollections, and does not claim to be comprehensive - there are undoubtedly many aspects of life at EISCAT that have been forgotten or passed over. To fill in some of the holes, I have relied on already available published information. EISCAT's Annual Report series, which provides continuing coverage of the system developments and the scientific production from the start, and which is now available on the EISCAT website www.eiscat.se, has been particularly helpful. Interested readers looking for more detail are encouraged to visit the website and do their own research.

The body of the paper is subdivided into 13 sections. Sections 1 to 6 describe how the Sites were organised according to the initial plans and how the organisation and staffing then gradually adapted to the actual demands, based on experience. They cover the work and responsibilities at the three sites, including important tasks not commonly known in the user community but vital to the ability to maintain operations and observations and deserving of recognition. The challenges involved in maintaining a geographically dispersed, multi-site high-tech system in the high North in the 1980s, with national borders, limited communications, no Internet, 1970s computer technology and a semi-Arctic climate, are described. The new tasks connected to the development of the Svalbard radar system are also mentioned, as are the Annual Review meetings and other social events.

65 scatter signals were coming on line and their performance was increasing by orders of magnitude every few years - so the technology was there, albeit large, costly and partly clumsy. Last but not least, the amount of sensitive electronic devices in ordinary households was very small or nonexistent, so the risk of a radar system established in a populated environment generating interference to consumer equipment was negligible.

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105 The paper is subdivided into this Introduction and a further 14 sections. To set the stage for the technical part, sections 2 to 7 are dedicated to the most important component of the system, the staff, without whom nothing would have been accomplished, and their working conditions. Section 2 describes how the Sites were organised and staffed according to the initial plans and how the organisation and staffing then gradually adapted to the actual demands, based on experience. Sections 3 and 4 cover the work and responsibilities at the three sites, including important tasks not commonly known in the user community but vital to the ability to maintain operations and observations and deserving of recognition. Section 5 addresses some of the challenges involved in

115 Section 7 deals with the organisation and work of EISCAT Headquarters (HQ) with special attention to the contributions of the Directors and the HQ software group.

Maintenance and repair tasks were largely defined by the radar hardware at the respective sites. Sections 8 and 9 therefore give overviews of the UHF and VHF systems. The prehistory and successes of both systems is summarized and their shortcomings are also discussed at some length, 120 in particular those of the VHF system, which are known to the user community only in very general terms.

A very important component in making EISCAT such a successful project was its computer system and the associated software system, EROS. Section 10 gives a description of these, how they were set up and operated and how experiment data was recorded and processed. 125

Another, possibly even more important component was the programmable correlator. This potentially very powerful but fault-prone and user-unfriendly device was eventually rehabilitated and its full potential unlocked, largely thanks to the work of a dedicated Site Programmer and an 130 equally dedicated Deputy Director. New coding schemes were then tested and implemented, first in Special Programs and after some time also in Common Programs, delivering data with much improved rate of statistics. Section 11 summarises these developments and presents a brief rundown of all Common Programs as they stood *circa* 1995.

135 Precise time and frequency keeping was essential to the success of the tristatic UHF system, but was non-trivial in the early 1980s - no satellite-borne navigation system open to the public existed yet - so had to be based on the use of atomic clocks at all sites and several backup systems. The maintenance and development of the timing system is briefly described in Section 12.

140 While the EISCAT system was established in the very sparsely populated far north of Fennoscandia, there was still a surrounding society with which it has had to co-exist on mutually acceptable terms. A critical aspect of this fact proved to be the spectrum issue, details of which are not widely known and therefore documented in Section 13 as a reminder to those possibly 145 contemplating the establishment of other active systems.

The paper ends with some personal reflections.

All work, results and successes documented here are products of a dedicated collective where everyone deserves equal credit. Apart from the Directors, individual staff members are therefore not 150 mentioned by name, except in a couple of cases where an important breakthrough can be ascribed to a single individual. External consultants, advisors and collaborators from the user community have been identified by name where relevant.

155 1. Site staff complement

The staffing and tasks of the three Sites in Finland, Norway and Sweden was defined in quite some detail already in the negotiation phase, subsequently laid down in the famous "Yellow Book" (Hagfors et al., 1974) and eventually formalized in the Agreement. In Finland and Sweden, 160 recruitment and employment of the site staff was to be subcontracted to the Kiruna Geophysical Institute (KGI, later IRF) and Sodankylä Geophysical Observatory (SGO) under a "matrix organisation" arrangement: site staff would belong to the line organisations of the local host

maintaining a geographically dispersed, multi-site high-tech system in the high North in the 1980s, with long distances, national borders, limited communications, no Internet, 1970s computer technology and a semi-Arctic climate. Section 6 briefly describes the new tasks connected to the development of the Svalbard radar system. Finally, Section 7 covers the Annual Review meetings and other social events aimed at fostering and maintaining a team spirit.

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institutions, but work full-time for EISCAT under direct control and supervision by Headquarters and be charged with operating and maintaining the radar equipment at their respective sites, thus effectively forming the operational branch of EISCAT. This scheme guaranteed the site staff the same benefits and social security as individuals in comparable positions in each host country, securing them credits towards their national pension plans and also offering a degree of job safety at the end of the planned lifetime of the Association (13 years). In Norway, the setup was initially different; in addition to being the Norwegian shareholder, the Norwegian research council, NAVF, would also employ the Tromsø site staff. This arrangement was gradually brought in line with that of the other two host countries, such that by 1995 the responsibilities for the staff were finally completely transferred to Tromsø University.

According to the Yellow Book, the total number of staff required at Tromsø during the first years was estimated to be 11, at Kiruna 5 and at Sodankylä 3 - but as the system was being constructed, these estimates were soon revised upwards. Initially, a Site Manager position was explicitly foreseen only for Tromsø, but was soon introduced also at the other sites. This position was the embodiment of "middle management". A skilled engineer by training, with both technical and management experience, the Site Manager was responsible toward HQ for all operational matters, including maintaining the equipment, executing the operational schedule and setting up and executing a site budget. At the same time, he was responsible toward the host institute for all personnel matters and toward his own staff for managing all day-to-day tasks like scheduling of shift work and looking after everybody's well-being. Previous management experience was probably decisive in the selection of the first group of Site Managers, who were all recruited from outside the host institutes.

Located at Ramfjordmoen, about 30 km south of Tromsø city and the university, the Tromsø site needed a range of skills that would enable it to function as a self-contained research station. To that end, service and support positions like a secretary, a caretaker and a combined mechanic/janitor were established. A substantial engineering staff, 8 - 9 positions, was going to be required for operating, maintaining and repairing the two transmitters and their accompanying antennas, receivers and signal processing systems. There was also a Site Scientist, responsible for the scheduling of the radar, the operation of the Common Programs and the support for visiting scientists coming to operate Special Programs. When the Heating system (Rietveld et al, 1990) was transferred from the Max-Planck-Institut für Aeronomie (Germany) to EISCAT in 1992, a dedicated Heating engineering position was also added to the Tromsø staff complement.

A much smaller staff complement had been planned for the remote sites, as these would be located relatively close to their host institutes and able to draw on service functions from these. Once in the operations phase, two engineers assisted by one technician would keep the sites running and operate the Common Program. In addition, they would handle system-wide maintenance, repair and improvement of the electronics subsystems developed at the respective site, viz. the receivers and timekeeping systems (Kiruna) and the ADCs and matched filters (Sodankylä). At Kiruna, the staff complement did indeed end up as foreseen, but the Sodankylä one was quickly augmented by two resident Site Scientists.

As the project got underway, it soon became clear that the amount of software and computer support required at the sites had been underestimated from the beginning, and each site eventually got a Site Programmer. That job title was really a bad misnomer, as the duties in practice did not only comprise coding but also - and perhaps more importantly - the maintenance of the site computer systems, including peripherals and operating system software; today a more proper job title would have been "software systems engineers".

160 2. Site staff complement

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2. Tasks during operations

- 215 During radar operations, work at Ramfjordmoen went on around the clock. Operation outside regular working hours (nights and weekends) required the engineering staff to work shifts. Since transmitter operation involved high voltages and possibly dangerous troubleshooting, there always had to be two persons scheduled per shift, one of whom had to be a competent transmitter operator. In addition to keeping the transmitter running, the receiver and data recording systems also had to be monitored and data tapes changed at regular intervals. Visiting scientists, coming to Tromsø for experiment campaigns, often volunteered to join the night shifts to off-load this duty from the staff. After a few such night shifts, they could return to their home institutions with an appreciation of the challenges involved in operating the system and the skill and commitment of the site staff.
- 220
- 225 Restarting the system after a massive crowbar (explained below) was a dreaded task during night shifts. The transmitter always went down to standby following a crowbar and had to be brought back up to full power gradually. Voltage transients caused by the current pulse through the spark gap made the lights flicker and often crashed both the correlator and the computer systems, necessitating a restart of the running experiment. If the crowbar had been triggered by a spontaneous klystron arc during the receive part of a radar cycle, the receiver front end amplifiers up in the antenna hubroom were very often damaged, forcing a total stop until repairs could be effected during normal working hours - getting at them required a climb up two ladders to the upper antenna platform, 20 m above ground!
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- 235 At the remote sites the operations-related workload was much less demanding. There was no transmitter to worry about and the amount of raw data recorded was much less than at Tromsø, so a data tape could last for a full 24 hours. During the first few years there was nevertheless a night watch kept, as the correlators frequently stopped or crashed, requiring a manual restart. But as the system gradually stabilised and the remote monitoring and control features of EROS began to be trusted, more and more of the nighttime operations were monitored from Tromsø. Local staff was then kept on call and could be alerted if needed.
- 240

3. Maintenance and repair work, hazards

- 245 The other side of the site work, one that most visitors only got superficially exposed to, was the maintenance and repair of the transmitters and antennas (addressed later). Transmitter work was performed under conditions largely similar to those encountered in heavy industry; dirty, heavy and partly dangerous. Troubleshooting and repairs often had to be performed under time pressure.
- 250
- The transmitter hall was a large industry-type sheetmetal building, housing two big oil-filled tanks for the klystrons and other high voltage components. There was also a large high voltage capacitor bank and the famous "crowbar". A travelling crane, spanning the full length of the hall, was used to extract the klystrons from the tanks when they or some related high voltage component had to be serviced. This was a delicate job; the VHF klystrons were about 5 meters tall and very heavy, but at the same time fragile and very intolerant of mechanical stress. Reinserting them in their sockets without breaking vacuum required precision and patience.
- 255
- The modulator decks with their specially processed switch tubes were also located in the tanks, fully immersed in the oil. A weak point in the transmitter systems, they frequently had to be hoisted, left to drip off and repaired or replaced.
- 260

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245 data tape could last for a full 24 hours. During the first few years there was nevertheless a night watch kept, as the correlators frequently stopped or crashed, requiring a manual restart. But as the system gradually stabilised and the remote monitoring and control features of EROS began to be trusted, more and more of the nighttime operations were monitored from Tromsø. Local staff was then kept on call and could be alerted if needed.

250

4. Maintenance and repair work, hazards

The other side of the site work, one that most visitors only got superficially exposed to, was the
255 maintenance and repair of the transmitters and antennas (addressed later). Transmitter work was performed under conditions largely similar to those encountered in heavy industry; dirty, heavy and partly dangerous. Troubleshooting and repairs often had to be performed under time pressure.

265 While the klystrons and the other equipment in the oil tanks constituted work hazards mainly because of their bulk and mass, the potentially most dangerous, even lethal, component in the transmitter hall was the capacitor bank (Figure 1, left). Its purpose was to supply the current for the klystron beam when the transmitter pulsed, while at the same time maintaining the beam voltage. It was constructed in two sections, one 80 uF and one 20 uF, that could be paralleled if required and charged to over 100 kV. When fully charged, it contained about a quarter of a kilowatt-hour or the equivalent of about 200 grams of TNT. Since an uncontrolled discharge inside the tube could dump all this energy into a plasma arc and instantly ruin the klystron, there was a protective mechanism, the "crowbar" (Figure 1, right), installed. This was a triggered spark gap, connected across the cap-bank, that was fired if the monitoring systems detected a rapid rise in the klystron beam current. A massive spark then formed in a couple of microseconds, effectively short-circuiting the cap-bank and dropping the klystron beam voltage to only a few hundred volts, thus eliminating any possibility of a tube arc - but the spark also generated an almighty bang that could be heard throughout the site. As a safety precaution, the cap-bank and the crowbar were enclosed in a netting cage with interlocked gates that would break the high voltage and trigger the crowbar if opened while the system was operating.



Figure 1: Left, an overview over part of the capacitor bank. Right, the "crowbar" spark gap. Photo courtesy of Ralf Larsen, Tromsø, formerly with EISCAT.

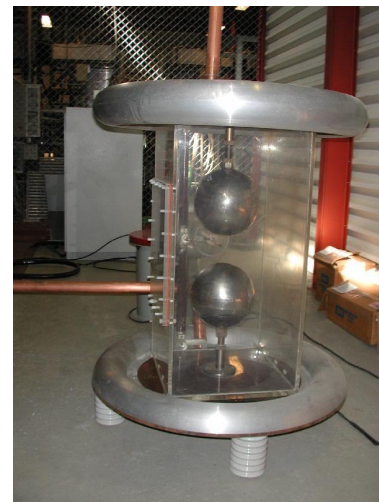
285 Another, invisible occupational hazard was X-ray radiation. When the transmitters were operating, the klystron beam collectors emitted X-rays with a maximum energy of 80 - 100 keV; most of the radiation was however the result of multiple scattering and therefore of much lower energy. The collectors were enclosed in lead shields, but even so some radiation leaked out and the engineering staff had to take care not to expose themselves unnecessarily. Everyone working around the transmitters wore film dosimeters that were checked at regular intervals by the radiation safety department at the university.

295 A frustrating aspect of the maintenance work was the fact that the original transmitter systems, as well as the 32 m UHF antennas, were designed and constructed according to the SAE engineering standards dominating in the US. Pump motors, including the big motor for the main coolant pump, were designed for US voltages and required extra transformers to adapt them to the Norwegian power system. Nuts and bolts throughout the transmitters had UNC or UNF threads and were specified to the nearest 1/64th of an inch, requiring SAE socket wrenches instead of the metric ones used on all European-made equipment. Pipe sizes were likewise specified in inches, keeping the site

260 The transmitter hall was a large industry-type sheetmetal building, housing two big oil-filled tanks for the klystrons and other high voltage components. There was also a large high voltage capacitor bank and the famous "crowbar". A travelling crane, spanning the full length of the hall, was used to extract the klystrons from the tanks when they or some related high voltage component had to be serviced. This was a delicate job; the VHF klystrons were about 5 meters tall and very heavy, but at
265 the same time fragile and very intolerant of mechanical stress. Reinserting them in their sockets without breaking vacuum required precision and patience.

The modulator decks with their specially processed switch tubes were also located in the tanks, fully immersed in the oil. A weak point in the transmitter systems, they frequently had to be
270 hoisted, left to drip off and repaired or replaced.

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mechanic busy with devising ad hoc adapters when the main coolant loop plumbing developed a leak - which it did from time to time with sometimes spectacular results. But when the cast-iron casing of the original main coolant pump developed a crack, the whole pump was replaced with a European made unit. From then on, as components began to fail, they were replaced with functional equivalents conforming to European standards wherever possible, which eventually simplified the maintenance task a lot.

4. Travel on the job

Maintaining the tristatic UHF facility required a lot of travel between the sites. During the buildup phase and the first three years of operation, the only road connection from Kiruna to Tromsø was via a narrow two-lane road through Finland, from Karesuando to Kilpisjärvi (see Figure 2), a distance of about 410 km, for the most part speed-limited to 80 km/h and initially involving a crossing of the Muonio river by road ferry at the Swedish-Finnish border at Karesuando; the bridge there was only built in 1980. This was shortened to about 330 km when the new E10 road from Kiruna to Narvik was opened in September 1984. Sodankylä to Kiruna is about 330 km and Sodankylä to Tromsø is almost 480 km. These distances made it impractical to visit one of the other sites for business and return the same day. Most trips tended to become at least two day affairs, except during the summer months when there was daylight around the clock - but all sites had facilities for overnighiting. During experiment campaigns these were also heavily used by visiting scientists. The Kiruna site building had a guest room that could house two, visitors to Sodankylä could use the Geophysical Observatory guest rooms, and in Tromsø a prefabricated barracks-type building, affectionally called the "Hilton" and containing eight guest rooms, a kitchen and a common room area was erected within walking distance from the site.

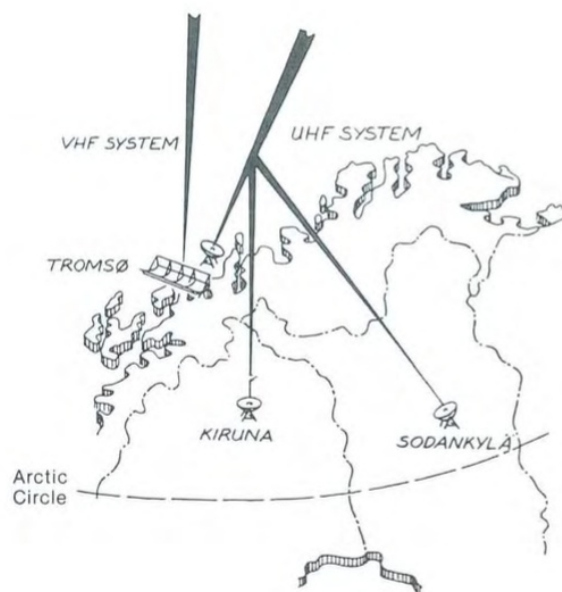


Figure 2: Outline map of the EISCAT KST geographical area (from the 1984 EISCAT Annual Report). The initial road connection between the Tromsø site and the remote sites entered Finland at Kilpisjärvi, close to Treriksröset, the point where the Swedish, Finnish and Norwegian borders meet. From there it continued southward on the Finnish side of the Swedish-Finnish border. The connection to Kiruna branched off and crossed the border at Karesuando, about 110 km from Treriksröset, and the road to Sodankylä branched off at Muonio, 80 km farther on.

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In the wintertime, driving was sometimes downright hazardous, most of the distance having to be covered in darkness and often with drifting snow reducing the visibility to a few meters. There was also a constant risk of encountering reindeer in the road. No staff were ever involved in serious accidents, but breakdowns occurred now and then and could result in long delays. My first trip to the Sodankylä site in mid-winter 1982 is a case in point - a front wheel bearing breakdown in Pajala, about halfway, almost ended my trip in the ditch but thankfully only resulted in an unplanned visit to the only garage in town and a total travel time of 18 hours for an average speed of 18.3 km/h...

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5. Development work for the ESR project

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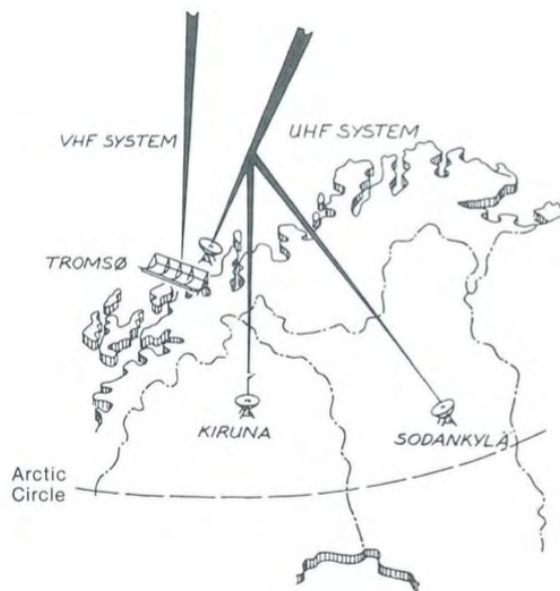


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6. Annual Review meetings and Christmas parties

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opportunity to get to know their colleagues at the other sites and learn about their work first-hand.
To help building and strengthening a team spirit, a Review Meeting of the whole staff was therefore
held every year. This was a three, sometimes four day combined work and socializing event, at
which everybody had an opportunity to present their work in conference-format sessions and
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of time for skiing, fun and play. The venue rotated between Finland, Norway and Sweden in a three
year cycle. Local arrangements were as far as possible handled by the staff of the host country site,
with help from HQ as needed. During the first decade the meeting was always arranged during the
best skiing season in mid-March, when the sun had returned and the snow still lay meter-deep, at a
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local restaurant, followed by relaxed socializing by the fireside. Spouses were also invited.

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7. Headquarters

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Figure 3: Group photo of the EISCAT staff at the 1985 Annual Review meeting in Abisko, Sweden. Most of those pictured here have long since left the Association and/or are now retired, but one individual is still working for EISCAT! From the 1985 EISCAT Annual Report.

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"EISCAT CEO". He was appointed by the Association's governing body, the Council, and
responsible for implementing the scientific program as defined by Council and committees. He was
also charged with maintaining relations with the local bodies in Norway, Sweden and Finland
hosting the EISCAT sites. His staff included an Assistant Director Science (ADS), responsible for
the scientific operations, particularly the Common Programs, and reporting to the Scientific
440 Advisory Committee (SAC), an Assistant Director Technique (ADT), responsible for the technical
infrastructure, a Business Manager, handling the Association's financial and business operations
and communicating with the Administrative and Finance Committee (AFC), and a secretary. One
or two administrative assistants (the number varied over time) were also employed. A very
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logical move, since the EISCAT system was the first of its kind and no comparable installation
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up to date on the progress, a very ambitious EISCAT Newsletter was launched under the editorship
of the ADS. Unfortunately, it seems that the workload at HQ soon overwhelmed the staff and the
Newsletter just stopped. Very few copies are known to remain. I have a copy of No. 3, which is full
of great information about the radar system and complementary instruments, like the STARE radar
and the GEOS II and ISIS-2 satellites. It was probably the last issue.

470 I had the opportunity to work under the Association's first four Directors, whose terms span the
scope of the present paper. Looking back, it is clear that each of them was uniquely equipped to
handle the tasks that dominated their respective time in office. As the first Director, Tor Hagfors led
the build-up phase. It would have been impossible to find a better qualified candidate. Hagfors was
a brilliant theoretician and one of a small group of people who had derived the theory of incoherent
475 scatter independently of each other. He also had extensive experience in science administration;
before starting with EISCAT in 1975, he had served as the director of the Jicamarca and Arecibo
radar observatories.

480 In 1982 Hagfors left for new challenges in the US and was succeeded by Murray Baron, a radar
scientist from SRI International in Menlo Park, California. Baron had a long career in ionospheric
radar and had been deeply involved in the design, implementation and exploitation of the
Chatanika, Alaska ISR system. At this point in time, HQ was working on solving the VHF
transmitter issue; Aydin, the California-based company contracted to build the transmitter was in

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485 difficulties and lagging seriously behind the delivery schedule (see Sect. 9). Coming from a workplace located in the same high-tech area and with an intimate knowledge of US law and business practice, Baron was the ideal person to resolve the issue. Under his guidance the transmitter was finally delivered in 1984 and all remaining business dealings with Aydin terminated.

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525 8. The UHF system

Looking back through the earliest documentation, it is obvious that the Steering Group was convinced of the feasibility of the UHF system very early on through contacts to industry and already operating ISR systems. The Yellow Book contains a highly technical description of the proposed transmitter, based on a feasibility study for the St. Santin radar and complemented by a detailed costing; the actual transmitter was largely patterned on this proposal. A detailed description of the antenna and receiver system planned at the time is also presented there, but this was of course superseded by the later decision to go for fully steerable antennas at all sites.

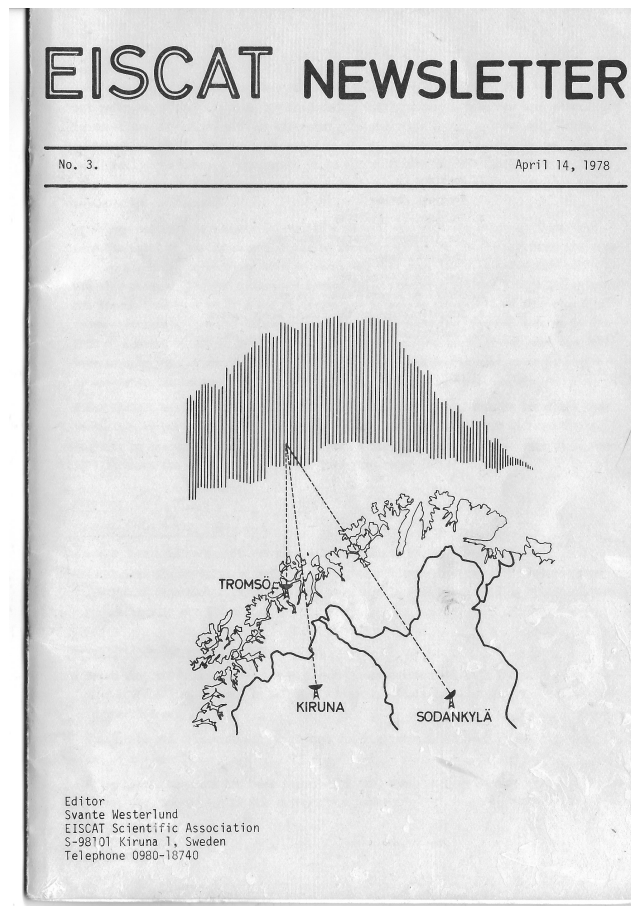


Figure 4: Front page of the EISCAT Newsletter, No. 3, issued by HQ in April 1978. Scan of my personal copy.

I had the opportunity to work under the Association's first four Directors, whose terms span the scope of the present paper. Looking back, it is clear that each of them was uniquely equipped to handle the tasks that dominated their respective time in office. As the first Director, Tor Hagfors led the build-up phase. It would have been impossible to find a better qualified candidate. Hagfors was a brilliant theoretician and one of a small group of people who had derived the theory of incoherent scatter independently of each other. He also had extensive experience in science administration; before starting with EISCAT in 1975, he had served as the director of the Jicamarca and Arecibo radar observatories.

In 1982 Hagfors left for new challenges in the US and was succeeded by Murray Baron, a radar scientist from SRI International in Menlo Park, California. Baron had a long career in ionospheric radar and had been deeply involved in the design, implementation and exploitation of the Chatanika, Alaska ISR system. At this point in time, HQ was working on solving the VHF transmitter issue; Aydin, the California-based company contracted to build the transmitter was in difficulties and lagging seriously behind the delivery schedule (see Sect. 9). Coming from a workplace located in the same high-tech area and with an intimate knowledge of US law and business practice, Baron was the ideal person to resolve the issue. Under his guidance the VHF transmitter was finally delivered in 1984 and all remaining business dealings with Aydin terminated.

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evolved to a point where the hardware was regarded as predictable and reliable. It was planned as a
two klystron system, but initially only one tube was installed. It was then run with just one klystron
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that good velocity estimates could be had using a common $\approx 350 \mu\text{s}$ pulse for all sites, so the long-
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much of the time. The cosmic noise background would not be a problem - at 933 MHz the sky noise
temperature is typically 10 - 15 K - but in the receiver, the signals would have to compete with
noise generated in the first amplifier stage, typically 10...100 times stronger than the signal,
555 resulting in signal-to-noise ratios of only a few percent. A prime design target for the UHF receiver
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Figure 3: The Kiruna EISCAT UHF antenna. The Tromsø and Sodankylä antennas were identical, apart from the feed system of the Tromsø one, that included a waveguide run with rotary joints connecting to the transmitter. Photo courtesy of Lars-Göran Vanhainen, Swedish Institute of Space Physics (IRF).

In the Intelsat application for which the antennas were originally designed, fast movement was not needed, as the antennas were used as the ground endpoints of links to geostationary communications satellites and therefore only needed to move slowly to optimize the pointing to the satellite or to move to a stow position. The EISCAT application was quite different; an ability to move rapidly was essential for many observations and so the drive motors and gearboxes fitted to the EISCAT antennas were much more powerful than those of the ancestor antennas. It soon became obvious that this mode of operation also resulted in a much increased need for regular service of the drives and the antenna structures.

Lubrication and routine maintenance of the drive systems was carried out by the local site staff. They also looked after the "pintle bearing" that kept the antenna centered on the rails. A number of plastic-coated segments, mounted on a ≈ 2 m diameter concrete cylinder centered on the azimuth axis, made up the fixed part of this bearing; the moving part consisted of four pads bolted to the alidade. The fixed pads wore out quite fast and had to be replaced about once a year on average, but soon an efficient routine evolved; worn-out pads were sent to Finland and reconditioned there by a company specialising in industrial plastics.

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It was well understood that the UHF signals backscattered from the ionosphere would be very weak much of the time. The cosmic noise background would not be a problem - at 933 MHz the sky noise temperature is typically 10 - 15 K - but in the receiver, the signals would have to compete with noise generated in the first amplifier stage, typically 10...100 times stronger than the signal, resulting in signal-to-noise ratios of only a few percent. A prime design target for the UHF receiver was therefore lowest possible noise temperature. It was clear that to reach the remote station target system noise temperature of $\approx 30 \text{ K}$, a cooled receiver front end would be mandatory. On advice from the radio astronomy community, the designers chose a solution based on a wideband, high gain, cryogenically cooled parametric amplifier, designed and built by AIL of Long Island, USA. This device could deliver about 60 dB of gain over a 30 MHz wide band while adding less than 20 K to the overall system noise. Unfortunately, it also turned out to be extremely sensitive to overload and ill suited for use in a radar system. At Ramfjordmoen, klystron arcs resulted in several burnt-out amplifier upconverters. These could not be repaired locally but had to be shipped back to AIL, causing major disruptions of the schedule. To maintain operation, the Tromsø system had to rely on uncooled transistor amplifiers with much poorer noise performance. These were eventually replaced by GaAsFET amplifiers designed and built at the Kiruna site, which brought the system noise down to 90-100 K. At the remote sites, the paramps were kept in operation for several years, but when the first cellphone base stations started up in Kiruna and Sodankylä their wide passband proved to be a liability; they went into saturation and had to be replaced by an in-house designed cooled GaAsFET amplifier system. The cryosystems were maintenance intensive and occupied a lot of the responsible engineer's worktime.

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Major maintenance work required calling in outside help. The alidade and the reflector backing were bolted together and the bolts required regular re-tensioning with a few years' intervals, a job requiring a crew certified for climbing and high altitude work - the top of the reflector backing was 35 m above ground! This job was therefore always contracted out. It was normally scheduled for a period of low demand in the summer and combined with a general overhaul of the reflector. To minimise the impact on the regular operation, all three UHF antennas were serviced one after the other in this manner during a single summer season whenever possible.

Bolt tensioning could be planned for in advance, but the same was not always true of the rail maintenance. After about a decade of operation, it was discovered at all three sites that the concrete carrying the rails had started to fracture in places and the rails were settling and even coming loose. This was a serious condition that had to be addressed promptly whenever detected, irrespective of season or weather: the antenna had to be locked in azimuth, the fracturing concrete chiseled away from underneath the rail and new frost-proof, rapid-curing concrete poured in, while monitoring that the section under repair ended up level with the rest of the rail. This work was always performed by contractors, but closely monitored by the site staff to ensure correct rail alignment.

For some time, the safety of the Finnish UHF antenna, located at the Tähtelä observatory about 10 km south of Sodankylä town, was in question due to actions by entities outside EISCAT control. This antenna stands on sandy soil just 80 m from the east bank of the Kitinen river. When the Association was established, the river and its tributaries were still unregulated, but in 1988 Kemijoki OY, the company possessing the water rights, applied for a permit to construct a hydropower plant at Kurkiaska, about 3 km downstream from the observatory. This involved damming the river and raising the water level by about six meters, so creating a water reservoir stretching tens of kilometers upstream from the dam. As soon as EISCAT got to know about the project, a strong statement of dissent was submitted to the water rights court; if the river were allowed to rise to the level requested in the project plan, the ground water table at the antenna would rise to almost two meters above the foundation footplate! This would introduce a risk for frost heave and in the worst case endanger the long-term stability and operability of the antenna.

Geotechnical experts were called in to consult both parties and long negotiations ensued. In the end Kemijoki OY agreed to a nearly three meters lower damming limit than first planned for, which ensured that the ground water table at the antenna would always be at least a meter below the bottom of the foundation. The company also agreed to put in insulation and a heating cable all around the foundation to protect against frost heave and to cover the operating costs of the heater for the lifetime of EISCAT. This agreement was made legally binding on Kemijoki OY by a decision in the water rights court on December 29, 1988. The heater and the insulation were put in the next summer. No problems related to the damming have been noticed since.

9. The VHF system

The 1971 Feasibility Study (the "Green Book") put forth strong arguments for a VHF system. Foremost among these was the ability to make measurements in conditions of very low electron density, in the mesosphere and the bottom of the E layer and in the topside ionosphere at altitudes well above 1000 km, where it was expected that it would detect the theoretically predicted "polar wind". The Green Book only contained a rough outline of the envisaged system, but the concept soon matured into a very ambitious design, comprising a 5 MW dual klystron transmitter, a 120 x 40 m parabolic trough antenna (Figure 4), steerable in the meridian plane between 30°N - 60°S, and a high-power RF switchyard, enabling operation in two distinct modes (Mode 1 and Mode 2).



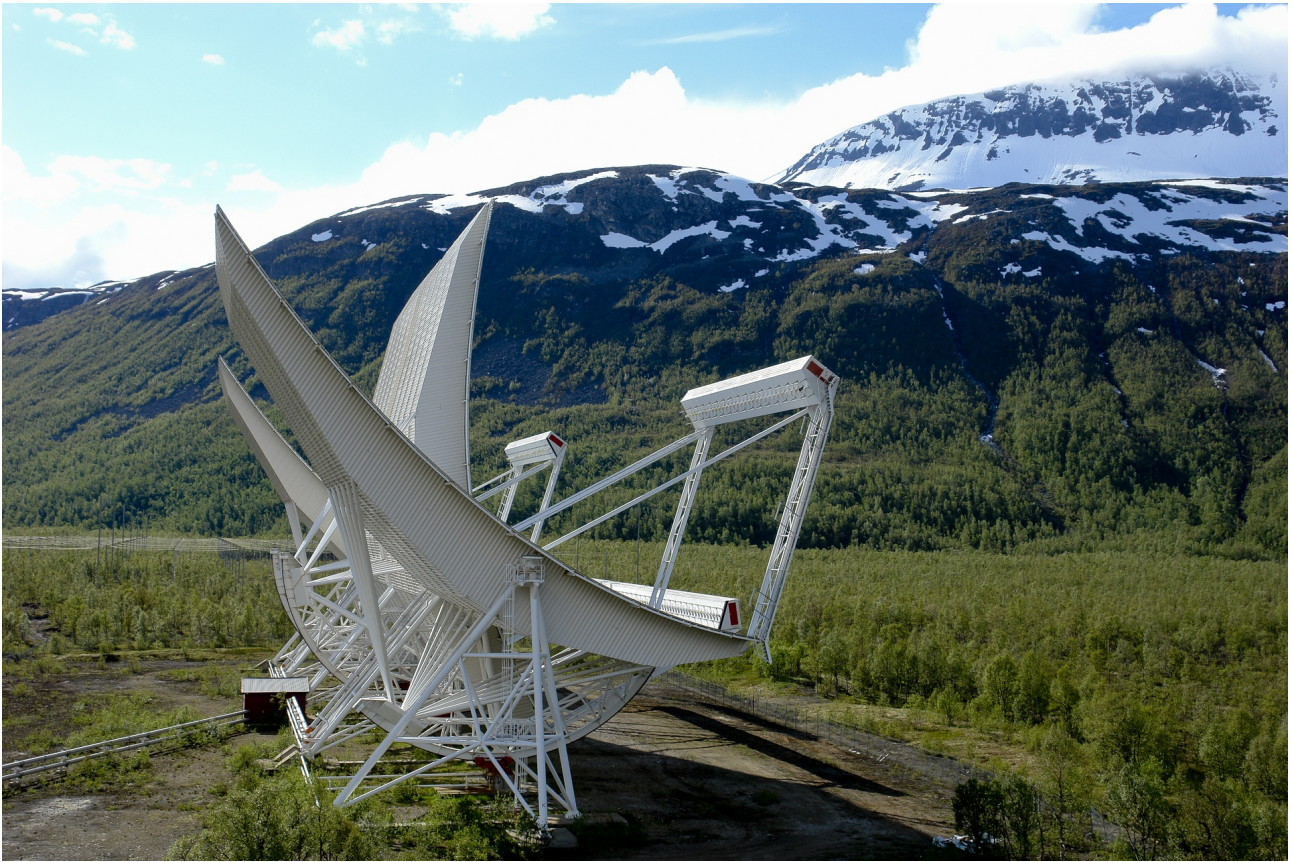
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665 **Figure 4:** The EISCAT VHF parabolic cylinder antenna at Tromsø. It is constructed as four individually tiltable 30x40 m reflector / feedline segments, which however must be electrically operated either as a single 120x40 meter aperture (Mode 1) or as two independent 60x40 meter apertures (Mode 2). Photo courtesy of Lars-Göran Vanhainen, Swedish Institute of Space Physics (IRF).

670 Mode 1 was optimised for maximum sensitivity. One klystron would drive all 128 horizontal dipoles in the antenna feed and the other klystron would drive all 128 vertical dipoles, making it possible to transmit either linearly or circularly polarized signals and to change the polarization on a pulse-by-pulse basis. With separate receiver chains for the two sets of dipoles, both total signal power and spectrum and Faraday rotation phase as functions of range could be recovered. In Mode 2, the antenna would be electrically split into two 60 x 40 m "half antennas", configured for circular polarization and fed by one klystron each. The two antenna halves could then be independently pointed in different directions both in elevation and in azimuth, effectively creating a dual beam system at the expense of sensitivity. However, for either mode to work as intended, there had to be two working klystrons in the transmitter, and precisely this turned out to be easier said than done...

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695 In parallel, EISCAT had started searching for an alternative klystron supplier. In 1983, the Valvo division of Philips was contracted to supply two klystrons with better specifications than the original Varian tubes. Delivery of the first Valvo tube, YK1320/1, was planned for December 1984 but slipped into early 1986, which allowed it to be immediately installed in the VHF transmitter; this had by then been operated at up to 2.5 MW with the only remaining Varian klystron.

700 The Valvo tube worked, but did not meet the performance requirements. It could be operated at 90 kV beam voltage, delivering about half of the contracted output power, 1.2 MW, but on raising the beam voltage further it started arcing; not even a 300 hour aging process improved this behaviour. Professor Tore Wessel-Berg of NTH and Jim Tallmadge of SRI International, both high power klystron experts, were called in to evaluate the situation in collaboration with the Valvo engineering team. After a thorough analysis, it was concluded that the tube design was basically sound but a number of modifications were required to make it operate stably at full power, all of which would require the tube be returned to Valvo for rebuilding. The proposed modifications could be directly incorporated into the second klystron, still under production at the Valvo plant. Left with no alternative, Council resolved to let Valvo modify the klystron according to the expert group recommendations.

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third was performing well below agreed specifications. The transmitter was finally delivered to Tromsø in 1984, and following this the contract with Aydin was terminated.

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With two functioning klystrons, the system could now be run in both modes. Mode 2 was put to good use in experiments designed for measuring the ionospheric convection velocity field over a wide area to the north of Tromsø. With the VHF antenna pointed to its lowest elevation, 30 °, and phased to generate two beams, one due north and the other approximately normal to the L-shells, Doppler velocities could be derived for a series of range gates along both beams and combined to generate two components of the 3D velocity. The third (semi-vertical) component could be provided by the UHF for the close-in ranges, but farther towards the north, it had to be extrapolated or was assumed close to zero. Variants of this experiment were used as Common Program 4. When the Svalbard radar came on the air in 1996, it was often run together with CP 4 experiments, pointing southward into a region where the fields of view of the two systems overlapped; under favourable conditions the full 3D velocity field could be determined over a large area in this way. Mode 1 suffered from a serious operational restriction: with the antenna pointing in the field-aligned direction and the transmitter operating at full power, radiation spillover from the upper edge of the antenna reflector was so strong that the Norwegian limits for public exposure to non-ionising radiation (which had been lowered since the antenna was designed) were exceeded even a kilometer to the south of the site. Four private homes along the main road suffered from bad interference to telephones, TV sets and audio equipment, and Tromsø site staff had to spend lots of time and effort on installing shielding, filtering and better TV antennas. A number of mitigation schemes were investigated, e.g. extending the main reflector, tilting the feeder bridge or erecting a 50 m high fence immediately south of the antenna, but all proved too complicated and costly. In the end, full power operation with the beam directed to the south of vertical had to be given up. This effectively eliminated one of the hoped-for core capabilities of the VHF system, regular high altitude field-aligned observations vital to the hunt for the "polar wind". Single-beam high altitude VHF operation thus became limited to vertical and remains so until today.

In the late 1980s, scientists at the Russian Polar Geophysical Institute (PGI) obtained funding for a receiving site for the EISCAT VHF transmissions to be built at Verkhnetolomsky on the Kola peninsula, about 510 km ESE from the Tromsø site. A large phased array, comprising eight 50 x 50 m 64-element modules and a matching dual polarisation receiver, handling both Mode 1 and Mode 2 transmissions, were to be constructed. The system was partially completed by 1991 and preliminary results were presented at the fifth EISCAT Workshop (Khudukon et al., 1990).

The PGI initiative was seen as a potentially valuable complement to the VHF system, so a number of tests were carried out, with the VHF transmitting long pulses vertically and the PGI system receiving. However, no signals were unambiguously identified. Time synchronisation remained a possible uncertainty, so a so-called TV sync receiver was prepared by the Kiruna staff and shipped to the PGI group. Unfortunately, this did not solve the problem. After a couple of years the PGI group dispersed and their fine system was left unfinished. Multistatic VHF observations thus had to wait until 2012, when the Kiruna and Sodankylä dishes were converted to VHF, after the introduction of UMTS mobile phone systems had made continued UHF operations impossible.

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10. Computers, computing and data handling

790 In the early 1970s, a European big science project many times larger and more complex than
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had finally got underway after years of financial negotiations, with the goal of constructing a 300
GeV accelerator at a total cost of up to 1150 million y1971 Swiss francs, more than 30 times the
projected y1974 cost of the EISCAT UHF system. To the dismay in some quarters and considerable
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been selected as the main control computer for the new accelerator, based on the proven reliability
of its predecessor, the Nord-1, and its real-time-gated architecture. At the request of CERN, Norsk
Data had provided the Nord-10 machines with a special input/output channel dedicated to
communicating with CAMAC, the de facto standard for fast nuclear and particle physics electronics
widely used at CERN. By 1976, a total of 25 Nord-10s were already in operation at different points
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All this was noted with interest by the EISCAT planners. Many of the processing demands of the
radar system, including the need for fast and predictable interrupt response, were similar to those of
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between HQ and the IRF EISCAT group and used both for software development and for data tape
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As delivered, the Nord-10 machines were equipped with 128 kByte of RAM. This was enough to
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put this in perspective, a USB memory stick can offer upwards of half a gigabyte of random-
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Tektronix monochrome unit displaying the autocorrelations computed from the raw data, using the
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Operator interaction with the EROS system was through ELAN commands, issued through a simple line-oriented user interface running on an alphanumeric terminal. This provided access to all major hardware, mechanisms to call up and run predefined experiments from file and the ability to interact asynchronously with a running experiment - possible because ELAN was an interpreted language. Below the user interface level, a large number of device-specific and mutually interacting programs and device drivers handled individual subsystems through CAMAC. EROS was quasi-realtime capable; the execution of commands could be scheduled to occur at specified points in time to a resolution of one second and there was a catch-up facility that enabled crashed and restarted experiments to seamlessly get back in sync with the other sites. Inter-site communication was arranged via leased telephone lines and made an integral part of EROS; it was possible to access the systems at the other sites, send messages, check hardware and data taking status and even command the antennas and start and stop experiments. As the system stabilised, this functionality was gradually used to eliminate the need for all-night staffing of the remote sites during experiments; visiting scientists often came singly to Tromsø and ran their special experiment campaigns from there, relying on the remote command facility to handle the remotes and trusting them to behave - which they did most of the time.

While the design, development and evolution of the core EROS system and the data handling routines resided with the HQ software group, the development of driver software for most subsystems was assigned to the Site Programmers. The UHF antenna control routines, as well as an advanced system for antenna pointing calibration based on radio stars, were developed in Sodankylä. Kiruna developed drivers for the receivers and real time clocks and Tromsø managed the initial stages of radar controller and correlator software development in collaboration with individuals from the Nordlysobservatoriet and NTH Trondheim, notably Hans-Jørgen Alker.

The original EROS system was in more or less continual evolution for about 20 years, particularly with respect to the data collection and recording parts that had to keep up with the gradually increasing data volumes and the introduction of new storage media (fixed hard disks, Exabyte tape etc.), taking up a considerable fraction of all programmers' time for years into the operations phase. Unfortunately, very little documentation of all this development work exists in accessible form today. Two important Technical Reports documenting the interfaces between EROS and EISCAT users have survived, one teaching the user how to program the radar using ELAN (Armstrong, 1980) and another describing how to access the data (Farmer, 1980). This latter report shows that much thought went into defining recording formats simplifying end user access to the data, while at the same time adhering to established standards (ANSI X3.27 and British BS4732). As far as the inner workings of EROS were concerned, it appears that the programming staff relied on the Fortran source code being self-documenting and therefore no reports seem to have been produced; none were found in the EISCAT HQ major search and review of old documentation carried out in 2020.

While EROS and its component parts thus controlled most of the radar system, two critical subsystems were handled individually: the Radar Controller (RC) and the Correlator. The RC generated the signals controlling the transmitter, ADCs and Correlator during a radar cycle; the Correlator was an application-specific, pipelined device that processed the received data "on the

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TARLAN was the Transmit And Recive LANguage for the RC. It had a simple syntax - individual hardware devices were addressed by symbolic names, and the time when an operation was to be performed was given as the number of microseconds after the start of a cycle. All commands had to be in time sequence, and transmitter commands had to be issued in a specific order and with certain minimum time separation to ensure a stable output waveform. Formally correct TARLAN code was translated by the compiler into a bit-level file, which could be loaded into the RC. Whether the code made the radar do what the experimenter was hoping for then had to be checked in a test run...

The history and specifics of CORLAN are tightly interrelated with the Correlator hardware design and therefore treated in the Signal Processing section.

The end product of all EISCAT operations was the "raw data", the time-averaged autocorrelation data output by the correlator, complemented by a block of metadata describing the system status (transmitter power, frequency, antenna pointings etc.). For the first few years, the extremely limited hard disk space forced a solution where raw data dumps were continually written to tape during experiments. Most experiments dumped data every 10 seconds. In Tromsø, a large tape then lasted at least 15 - 16 hours if the experiment did not crash (a crash often led to a forced tape change and a restart), but experiments running over several days involved a number of tape changes, often during the night shifts. The data volume at the remotes was much smaller, as only some three signal-carrying range gates, centered on the beam intersection point, were computed.

Initially, it had been estimated that the Tromsø system would generate in the order of 100 data tapes per year, but as experiment operations picked up and new coding schemes were introduced, the number of tapes grew dramatically. As a result, tape handling expanded into an almost full-time job. EISCAT was responsible for making archive copies of all data tapes and archiving them securely. In addition, user copies had to be made and forwarded to the data representatives in the member countries. Copies of Special Program data collected during national experiments were only sent to the respective member country, whereas all Common Program data had to be distributed to all member countries, requiring six copies. The copying and archiving job was centralized to HQ in Kiruna, so tapes had to be physically transported there as soon as convenient after an experiment. As shipments from Tromsø and Sodankylä always involved a trip by car, a routine soon developed where tapes were accumulated locally until the shipment could be combined with a visit by a staff member for some other purpose. Sometimes two, three or more boxfuls of tape were shipped at once, creating a massive peak in the workload at HQ. Eventually a part-time Data Assistant position with responsibility for all tape handling had to be created.

After a few years, user demand and the availability of more computing power in the form of the ND-530 machines led to the introduction of a near-real-time quick-look analysis program, which produced first order estimates of standard parameters (plasma density, electron and ion temperatures and velocities) "on the fly". These estimates were also saved and distributed to users. Following the installation of a ND-5400 machine in 1990, EISCAT HQ was connected to the Internet via the IRF and its connections to SUNET, the Swedish university network. From this time onwards, data distribution was gradually moved from physical media to file transfers over the Net, eventually putting an end to all shipping of tapes to the Associates.

While EROS and its component parts thus controlled most of the radar system, two critical subsystems were handled individually: the Radar Controller (RC) and the Correlator. The RC generated the signals controlling the transmitter, ADCs and Correlator during a radar cycle; the Correlator was an application-specific, pipelined device that processed the received data "on the fly". Both were operating at microsecond- or sub-microsecond time resolution, had to be programmed at the bit level and were best left alone once started and running. For these reasons, two unique languages, TARLAN and CORLAN, were constructed. Both were compiled - the logical design choice for languages constructed to control units with the ability to cause unpredictable and even dangerous system behaviour if handled randomly.

TARLAN was the Transmit And Receive LANguage for the RC. It had a simple syntax - individual hardware devices were addressed by symbolic names, and the time when an operation was to be performed was given as the number of microseconds after the start of a cycle. All commands had to be in time sequence, and transmitter commands had to be issued in a specific order and with certain minimum time separation to ensure a stable output waveform. Formally correct TARLAN code was translated by the compiler into a bit-level file, which could be loaded into the RC. Whether the code made the radar do what the experimenter was hoping for then had to be checked in a test run...

The history and specifics of CORLAN are tightly interrelated with the Correlator hardware design and therefore treated in the Signal Processing section.

11.3 Data handling

The end product of all EISCAT operations was the "raw data", the time-averaged autocorrelation data output by the correlator, complemented by metadata describing the system status (transmitter power, frequency, antenna pointings etc.). For the first few years, the extremely limited hard disk space forced a solution where raw data dumps were continually written to tape during experiments. Most experiments dumped data every 10 seconds. In Tromsø, a large tape then lasted at least 15 - 16 hours if the experiment did not crash (a crash often led to a forced tape change and a restart), but experiments running over several days involved a number of tape changes, often during the night shifts. The data volume at the remotes was much smaller, as only some three signal-carrying range gates, centered on the beam intersection point, were computed.

Initially, it had been estimated that the Tromsø system would generate in the order of 100 data tapes per year, but as experiment operations picked up and new coding schemes were introduced, the number of tapes grew dramatically. As a result, tape handling expanded into an almost full-time job. EISCAT was responsible for making archive copies of all data tapes and archiving them securely. In addition, user copies had to be made and forwarded to the data representatives in the member countries. Copies of Special Program data collected during national experiments were only sent to the respective member country, whereas all Common Program data had to be distributed to all member countries, requiring six copies. The copying and archiving job was centralized to HQ in Kiruna, so tapes had to be physically transported there as soon as convenient after an experiment. As shipments from Tromsø and Sodankylä always involved a trip by car, a routine soon developed where tapes were accumulated locally until the shipment could be combined with a visit by a staff member for some other purpose. Sometimes two, three or more boxfuls of tape were shipped at once, creating a massive peak in the workload at HQ. Eventually a part-time Data Assistant position with responsibility for all tape handling had to be created.

970

11. Signal processing

The ultimate purpose of an incoherent scatter radar observing the ionosphere is to determine the physical state of the scattering plasma: its electron and ion temperatures, ion composition and bulk velocity. Any particular combination of these parameters results in a corresponding distribution of scatterer (electron) velocities, which in turn manifests itself as a specific power frequency distribution of the scattered signal. After processing by the radar receiver, the signal is sampled at regular time intervals and digitised, such that the output from the ADC forms a discrete-time series of signal complex amplitude estimates. A straightforward way to extract the desired information from the signal is now to compute time-averaged autocorrelation functions (ACFs) over segments of the sample stream. When EISCAT was in the planning stage, this time-domain approach was already used at the Chatanika radar, where two different digital correlators were in use, one hardwired for long pulse modulations and the other semi-programmable (by rewiring) for multipulse modulations. It was decided to follow this route also in EISCAT by employing the famous "Alker correlator".

The Alker correlator promised to be a major improvement over the Chatanika units. It started as a doctorate thesis project at the Norwegian technical university in Trondheim (NTH), with design targets set by the envisaged requirements of the coming EISCAT system and (at least initially) supervised by Tor Hagfors. The result was a software programmable device optimised for computing complex-valued correlation functions, potentially very flexible and in principle able to handle experiments combining several different types of modulation in each radar cycle. A detailed description of its pipelined architecture is given in (Alker, 1979). The design used the latest state-of-the-art components, some of which were not even commercially available as the work started: AMD 2900 bit-slice processors, TRW LSI 8x8 bit multiplier chips and static NMOS chips making up the input and output memory banks. The program flow could be controlled by conditional jumping, based on the values of three programmable loop counters, which allowed a degree of structured programming. Unfortunately, the program memory was only 63 instructions deep (it had to be very fast and was therefore expensive), which was a severe limitation on an otherwise beautifully thought-out design.

The first physical realisation of the correlator turned out to be marginal. It was constructed as two separate 19 inch rack-mounted units interconnected by flatcables, one containing the double-banked input memory and the other containing the arithmetic unit (the only part of that unit actually constructed on printed circuit boards) and all control logic. The signal ground provided by the flatcable connection between the two units proved to be too weak for the 5 MHz data rate. The ADCs used 2's complement data coding, so signal voltage variations around zero volts caused all data bits on the bus to change from all zeros to all ones or vice versa. Because of the weak ground, this created lots of common-mode digital noise on the bus which partly corrupted the desired signal.

Internally, both the buffer memory unit and the computing unit were constructed using wire-wrapping, a construction method extensively used for logic systems in the 1970s and 80s which has now largely disappeared. It was useful for producing prototypes or small series of non-standard systems, but when improperly applied it could produce all sorts of illogical malfunctions. Unfortunately, this was the case with the correlators. During the first year of operation, the Kiruna unit could now and then stop in the middle of an experiment and refuse to load and start, leaving the site staff with no option other than to start pulling circuit boards. After a while, we developed a very direct troubleshooting technique: the circuit boards were banged down on the kitchen table, wire-wrap side down. This often caused one or several pieces of overstretched and broken-off wrap wire to fall out of the wire mat; it then remained to locate the broken connection and put in a new wire!

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four correlators (three for the UHF system + one for the VHF) and the demand for experiment
995 operations was increasing all the time, leaving no time for preventive maintenance. But in January
1983, a capacitor in the Kiruna correlator exploded and set fire to some internal cabling, spreading
residue inside the unit and releasing acrid smoke that triggered the fire extinguishing system. I was
there when it happened... Fortunately, all equipment was insured, and a contact to the insurance
company confirmed that the repair costs would be reimbursed in full. A rehabilitation program was
1000 now started. All electronics in the receiver room was thoroughly cleaned to prevent corrosion from
the hydrochloric acid fumes. At the same time, the VHF correlator was transferred from Tromsø to
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process, new result memory boards were constructed, doubling the result memory space to 4096
addresses. During 1986, a further improvement was introduced in the form of an internal 2x16k
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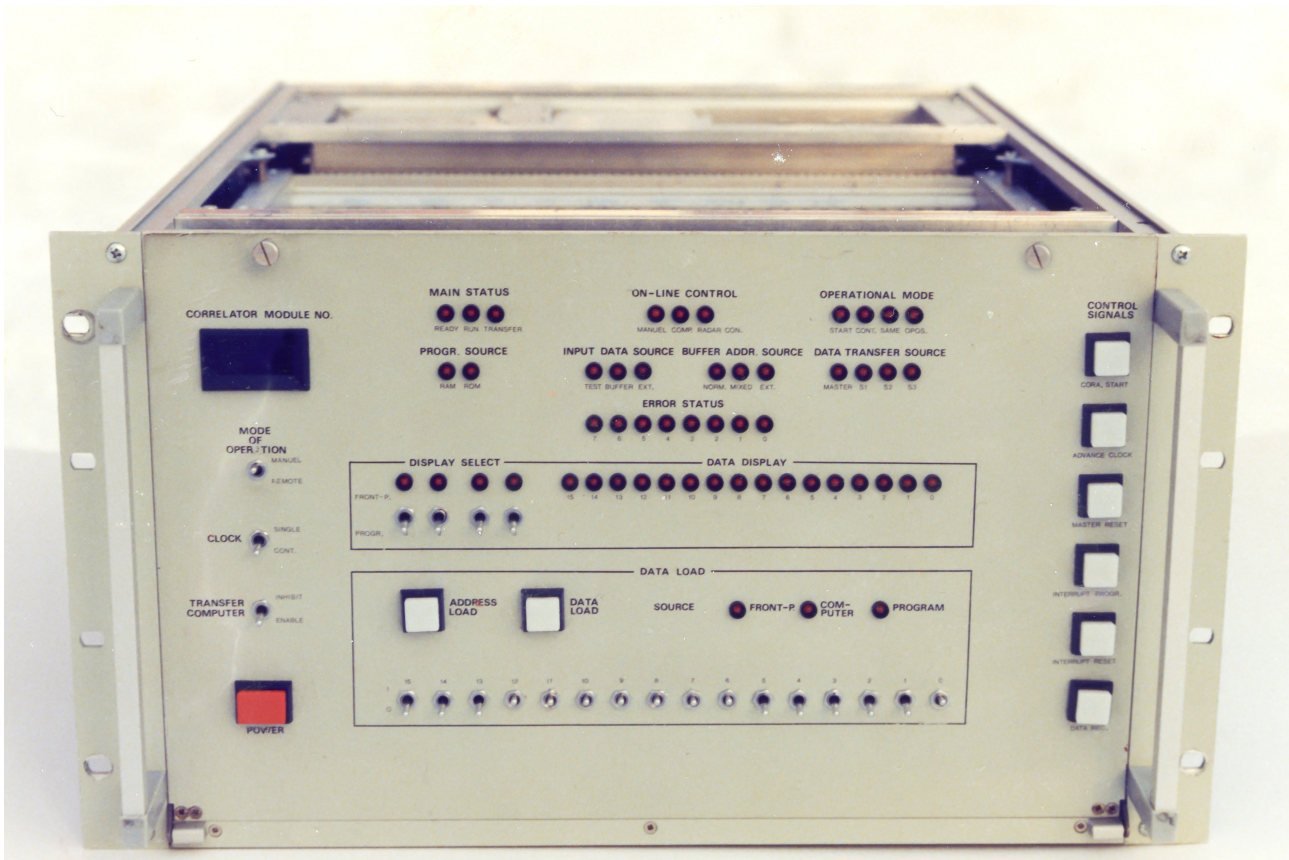


Figure 7: The arithmetic unit of one of the "Alker correlators". The front panel layout has many features reminiscent of minicomputers of the 1970s era, in particular the "data load" field with 16 switches to set up and manually load individual micro-instructions into arbitrary program memory locations; the front panel of the Nord-10 computer had an almost identical field for the same purpose. Photo courtesy of Lars-Göran Vanhainen.

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12.2 Correlation software - CORLAN and Uniprolog

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1060 For some time it now appeared that the radar was being used almost to its statistical limit - and this was indeed true as far as the "standard" modulations went. But in parallel to Turunen's work, different forms of power-domain coding had been studied by several authors and found promising, particularly in high spatial resolution, low signal-to-noise ratio situations like e.g. in the E layer at night, where they were expected to outperform multipulse group codes by 2 - 4 times or more. Perhaps the most well-known were the so-called *alternating codes* (Lehtinen and Häggström, 1987), which were derived from *Walsh functions* (e.g. Wikipedia).

1070 Lehtinen, Häggström, Vallinkoski and coworkers tried alternating codes on the UHF systems by using UNIPROG to compute all lagged products of the received samples from each code, dumping the intermediate, undecoded results at brief intervals and performing the decoding off-line after the experiment. This approach worked and generated high quality data, but it also suffered from severe limitations, perhaps the worst of which was the very restricted number of ranges that could be handled and the necessity to use only 8 bit alternating codes; the limited result memory precluded using the longer codes.

1075 It was clear that once the lessons from these first alternating codes experiments had been absorbed by the user community, one could expect a demand for the provision of user-friendly alternating codes capabilities at both the UHF and the VHF, possibly followed by an upgrade of the Common Programs. It was equally clear that the correlators were incapable of handling this job as they stood. In simple terms, decoding an alternating codes experiment reduces to computing a set of polynomials, where each term in each polynomial is the product of an accumulated lagged product, fetched from the result memory, and a sign bit generated from the code set. But when the correlators were designed in the late 1960s, the need for this type of operation was not anticipated and so the only arithmetic function implemented on the result memory side was straightforward accumulation.

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We were thus faced with a choice: either replacing the correlators with high performance workstations, which were only then becoming available, or accepting the challenge of somehow modifying the correlators to enable on-the-fly decoding of alternating codes, while leaving all other functions intact. After careful weighing of the pros and cons, the second alternative won out; it would allow UNIPROG- and GEN-system based experiments to run as before, almost transparently, while at the same time avoiding the uncertainties, development work and extra costs associated with porting the signal processing task to new, unfamiliar hard- and software systems.

After quite a bit of thinking, a solution to the decoding problem was found, using a previously unused data port into the correlator arithmetic unit to interface a newly designed sign bit generator/multiplier unit between the buffer memory and the correlator arithmetic unit. The concept was worked out, prototyped and tested in the Kiruna correlator. Once found working as planned, the added functionality was made accessible to users through a special driver, ALTCODE, which was added to the GEN-system library. Decoders were eventually installed in all correlators and a "second generation" programming environment, G2, combining the GEN-system and the alternating code capability, was developed and published (Wannberg 1993). G2-experiments were first used in regular scientific experiments in late 1990 with good results. Next, CP-1K, a new version of CP-1, was developed, where the interlaced multipulses of the previous version were replaced by a 16 baud alternating code modulation. The same modulation pattern was also soon introduced into CP-2, yielding considerably improved performance in low signal-to-noise conditions. After these developments, the signal processing hardware was essentially left untouched and performed well for a decade, until finally replaced by a system patterned on the one developed for the EISCAT Svalbard Radar.

Throughout the first two decades, the Common Programs continued to meet the basic concepts drawn up at the outset as far as their areas of coverage were concerned, in agreement with the idea to generate a data base covering as long a timespan as possible - but their performance increased all the time thanks to the improvements on the signal processing side. CPs 1, 2 and 3 were UHF experiments. CP-1 was a field-aligned experiment with altitude coverage to above 600 km, tristatic velocity measurements in the F region and km-scale altitude resolution in the E region, CP-2 was essentially the same experiment, but scanning through four closely spaced beam directions enclosing the Tromsø field line. CP-3 implemented an F-region latitude scan covering the entire common field of view of the UHF system. CP-4 was a derivative of the British Polar experiments, using the VHF in dual beam mode to measure the plasma velocity field to the north of Tromsø to Svalbard and beyond. CP-6 was used for high resolution measurements in the E and D regions, CP-7 was a dedicated high altitude VHF experiment. The modulations used were essentially variants of Turunen's GEN programs or (in the case of CP-1 and 2) copies of CP-1-K and remained basically unchanged until the signal processing system was replaced, starting around y2000.

12. Time and frequency keeping

Operating the tristatic UHF with pulsed transmissions, instead of with CW, required very precise relative time-keeping between the three sites. With the most probable pulse lengths being in the 300 - 400 us range, a relative clock drift of at most some 5 % or 15 - 20 us would be tolerable. This was doable, but not simple, in the 1970s - this was long before the time of GPS and other satellite-borne systems distributing reference time and frequency and so the only way to guarantee this level of accuracy over any reasonable length of time was by having Cs beam clocks at all sites.

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1140 A typical Cs beam clock is expected to drift by less than $1 \text{ e-}12$ relative to an ensemble of similar clocks. This translates to a little less than one-tenth of a microsecond per day, so with three well-behaved clocks one could expect the system to stay in time for about a hundred days or more, once the clocks had been set to a common reference, the "master clock". This was a fourth, battery-operated Cs clock that could be transported between the sites by car. Clock transports typically happened about once a year or whenever the signals received at Kiruna and/or Sodankylä started to drift out of the expected reception windows.

1145 The Cs clocks also generated an extremely stable 5 MHz frequency reference signal, which was used to phase-lock all oscillators in the transmitter and receiver systems. In this way the relative TX-to-RX frequency uncertainty at 933 MHz was less than $2 \text{ e-}3 \text{ Hz}$, negligible compared to the $\approx -5 \text{ Hz}$ residual shift of the transmitted frequency caused by klystron phase pushing during the pulse.

1150 However, the Cs clocks lacked a very important feature - while they generated an extremely accurate train of 1/s pulses, settable to better than a microsecond, they had no built-in machine-readable time output port. To remedy this, EISCAT staff designed and constructed a Real Time Clock (RTC). Controlled by the 1/s ticks from the Cs clock, this unit provided both a time display and a time output port. It also generated start pulses for various time-critical subsystems, primarily the radar controller unit. For this purpose it was fitted with a programmable delay register that could be loaded from the system computer.

1160 As an extra fall-back, each site also had a Loran-C receiver tuned to the Loran-C transmitter at Bø in Vesterålen, which transmitted an extremely precise, Cs clock controlled signal at 100 kHz. Special receivers comparing the local clocks against the sync pulses transmitted by the Yllästunturi TV transmitter were also used to monitor the relative drift of the Kiruna and Sodankylä clocks.

1165 To facilitate global experiment control and data analysis and archiving, the site clocks had to be synchronised to Universal Time (UTC). From time to time, the master clock was therefore transported to the Swedish national time and frequency standard, a pool of high-performance Cs clocks maintained at the Swedish defence research establishment (FOA) in Stockholm, and there reset to UTC to the nearest microsecond. The transport was normally effected by booking two tickets on a regular flight from Kiruna to Stockholm, one for the clock and one for the timing engineer, and seating the clock in its booked aircraft seat, a procedure which would be unlikely to be allowed now... As an amusing aside, on one occasion during the "technical period" the traveling clock was flown from Kiruna to Sodankylä in a small private seaplane piloted by the then Assistant Director Science, who held a private pilot license. The plane landed on the Kitinen river, next to the EISCAT site, only to be immediately inspected by two serious customs officials who had been ordered there for the occasion...

1175 Part way into the construction period it was realized that making all three UHF antennas steerable added a new degree of complexity to the timing system. Since the tristatic intersection point could now be anywhere in the common field of view of the three sites, the time between the transmission of a pulse from Tromsø and the arrival of the scattered signal at a remote site could be anything from about $600 \mu\text{s}$ to 20 ms - but the reception window at the remote sites could only be made about 1180 a millisecond long due to the limitations imposed by the correlator.

1185 An elegant combination of hardware and software resolved this issue. Whenever an antenna pointing command was executed at a remote site, the propagation time from Tromsø to the beam intersection point and thence to the receiver was automatically computed by an EROS operating system routine and loaded into the RTC delay register. The RTC then delayed the start pulse to the

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1190 **13. Frequency spectrum issues**

1195 Dictated by the plasma physics of the ionosphere, the optimal operating frequencies of incoherent scatter systems lie in the VHF - low UHF region, approximately 50 - 1000 MHz, the VHF frequencies being most suitable for low electron density conditions. To maximise the scientific returns from the EISCAT system over the whole 75 - 2000 km altitude range, the plan was therefore to obtain a UHF allocation just below 1000 MHz and a VHF allocation at approximately 240 MHz (du Castel et al., 1971). Also, to achieve the full potential of the project, access to up to 30 MHz of interference-free spectrum centered on the allocated operating frequencies was desired; only so would it be possible to employ frequency-hopping in order to use the full 12.5 % transmitter duty cycle and simultaneously receive both up- and downshifted plasma line returns.

1200 Unfortunately, most of the 50 - 1000 MHz range is allocated to a great number of active services (FM broadcasting, TV, cellphone systems etc.) and heavily congested everywhere, also in Scandinavia. The EISCAT allocations therefore had to be fitted into already established frequency plans on a mutual non-interference basis without deviating too much from the initial targets.

1210 In the late 1970s this was still possible. VHF TV channel 12 was not used in northern Norway, which left a convenient spectrum slot for a VHF system between 222.75 and 230.0 MHz. A transmitting permit for 224 +/- 2.5 MHz was applied for from the Norwegian P&T at an early stage and duly issued. This allocation proved to be fairly unproblematic; some interference, probably emanating from TV transmitters farther south, could be observed from time to time but did not seriously upset the VHF observations.

1215 For the UHF, an allocation at about 930 MHz was desirable, based on experience from the French St. Santin system. This also appeared possible, as the 918 - 948 MHz frequency band was not used by any radio services in Norway, nor in Sweden or Finland. Nevertheless, the band could not be allocated to EISCAT on a protected basis, because in the ITU Region 1 frequency plan first priority to its use was given to "fixed-to-mobile communications", that is, cellphone systems. In Sweden, a pure receiver system like the Kiruna UHF station did not require any special permit. On the other hand, Swedish telecom law did not provide any mechanism for granting interference protection to a receiver site. Rules in Finland were similar. But at this point in time, no collision of interests was foreseen; the EISCAT project was expected to last for about 13 years and terminate well before any mobile networks would be deployed in northern Scandinavia. On this assumption, a transmitting permit for 933.5 +/- 4 MHz was issued by the NPT, the UHF system was set up accordingly, and the community looked forward to many years of undisturbed wideband operation.

1225 But reality proved to be different, the initially promising spectrum situation soon developed into an existential threat to the whole UHF system. A new cellphone service using the 935 - 942.5 MHz band, NMT900, was introduced in the metropolitan areas of all Nordic countries already in December 1986. It rapidly expanded northwards and reached Kiruna in early 1988. The NMT900 base station signals immediately drove the Kiruna receiver into non-linearity, even though they were outside the EISCAT band. To rescue the continued operation, the UHF receiver front ends had to be completely redesigned and rebuilt to increase their dynamic range, the UHF transmitter frequency band was downshifted by 2 MHz (the limit of what the klystron could handle) and agreements were reached with the NMT operators in all three host countries to restrict the NMT900

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base stations near the EISCAT sites to transmit only above 939 MHz. It was only by August 1990 that these actions finally enabled the UHF to return to routine tristatic ion line operation, but plasma line observation possibilities had now become severely restricted in the process.

1240 However, the EISCAT interplanetary-scintillation (IPS) program was badly affected. Being a
passive technique, IPS requires access to wide segments of interference-free spectrum to obtain
statistics, a feature it shares with radio astronomy observations. At EISCAT, the IPS groups had
until then been able to use the full available UHF bandwidth, 30 MHz, but now became restricted to
8 MHz or less; further bandwidth cuts in the future could not be ruled out. An in-house
1245 development program was therefore started with the goal of establishing receiving capabilities in
the 1410-1427 MHz protected radio astronomy band at Kiruna and Sodankylä. When completed,
the 1400 MHz system performance matched the best previous 933 MHz results and gave the IPS
program an extra ten years.

1250 In an effort to obtain support for EISCAT's continued need for undisturbed spectrum, HQ contacted
CRAF, the Committee for protection of Radio Astronomy Frequencies, a sub-committee of the
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frequency bands and combating spectrum pollution. While protection of active radar operation did
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admitted with observer status, EISCAT eventually gained full membership. From the early 1990s
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1260 EISCAT in the radio astronomy community.

In 1998, the Association's initially planned-for 13 years of operation had come to an end, but the
system was generating very good science, and users and owners both wished to continue the
operation. This required an extension of the transmitting permits. Negotiations with the NPT
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this did not at first cause major interference problems, but in 2005 the Sodankylä UHF receiver was
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in Sodankylä town to transmit in the 929.0 - 935.0 MHz range, which was protected under an
agreement between EISCAT and the Finnish P&T. The operator was approached and rapidly re-
programmed all stations, which eliminated the worst interference. Unfortunately, transmissions in
this frequency range originating from other base stations more than 100 km away still affected
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the area, the useable Sodankylä receiving band shrank further and eventually became restricted to
just 929.0 - 931.5 MHz, a dramatic deterioration from the 1980s situation.

These problems notwithstanding, routine UHF ion line operations continued for more than another
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1295 **Personal reflections**

Today, more than 45 years after the inception of the EISCAT KST, its successor, EISCAT_3D, is faced with a whole new situation. Solid state technology has advanced to the point where computational power many orders of magnitude larger than that of the EISCAT correlators is available in every laptop, thus trivializing the signal processing task. But at the same time, the explosive growth of digital cellphone systems and the transition from analog to digital TV broadcasting has caused a massive demand for low UHF spectrum, leaving almost no holes below 1 GHz. Protected, undisturbed access to frequency bands wide enough to cover the whole scatter spectrum can no longer be had in developed areas like e.g. northern Scandinavia, and no help is to be expected from the radio astronomy community, which regards any transmission as anathema. It was only with difficulty that a slot at 233 MHz could be identified and accepted by the Norwegian spectrum management authority as suitable for the active part of EISCAT_3D. Neighbouring bands are filled with digital audio broadcasting signals, some of which will spill over into the radar band and increase the noise level there, which is likely to make plasma line reception hard or maybe even impossible. The proliferation of consumer electronics has also made the general interference situation much worse. One can only hope that advanced coding and signal processing can help out.

I fondly remember the simpler times of the 1980s and 90s and wish the new generation the best of luck in meeting today's and tomorrow's challenges.

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15. Personal reflections

My background and route into EISCAT were quite different from most of my colleagues. In 1979, on completing my Ph.D. in nuclear physics at Uppsala University, I started to look for a job. At about this time, EISCAT HQ was advertising in major Swedish newspapers. The EISCAT project looked very technically attractive and I felt that maybe it could use me somehow; I had been trained as an experimentalist, much of my work had been performed in an international environment at CERN, our system contained a great deal of radio frequency technology - and I was also an enthusiastic radio amateur who once dreamt of getting into radio astronomy professionally.

That spring, I applied for a Scientific Programmer position at HQ, was interviewed and turned down - but a few weeks later I got a phone call out of the blue from the Swedish Institute of Space Physics. They had got hold of my application to EISCAT and liked it, I was invited to an interview - and the next day I was offered a research engineer position in the satellite group! I jumped at the opportunity and started on the job in November 1979, spending the first couple of months studying up on plasma physics, but all the time I kept an eye out for any opportunities at EISCAT. When the Kiruna Site Manager position was about to become vacant in 1981, I applied for it and got it, perhaps because I was the only applicant - Professor Hultqvist regarded my choice as a voluntary demotion!

Taking up my new position in June 1981, I was still a green newcomer to the geophysics community and my understanding of the scientific tasks that the Founding Fathers of EISCAT and the Associate scientists had set for themselves was almost nil. In a way, the whole time from then until I departed EISCAT in 2008 came to be a continuous learning-on-the-job process, apparently with some success; in 1987 I was recruited by Jürgen Röttger to fill the HQ position of Assistant Director (Technique) and later promoted to Deputy Director, a position I held until my departure. My contributions to the development of the radar system peaked during the Svalbard radar project and culminated in my leading the EISCAT_3D feasibility study. I always regarded my role in EISCAT as that of a *machine physicist*, a concept I learned of at CERN, but I am happy that in helping to introduce the study of meteor head echoes with the UHF system I was also able to do a bit of science with the instrument. I am extremely grateful that I was able to join the EISCAT community right at the beginning of the operational phase and given the opportunity to help developing and steering the radar system through a period that was arguably its best years, when the VHF and UHF spectrum windows were still wide open, new and unexpected results were emerging regularly and coding and signal processing theory and technology made great advances.

Today, more than 45 years after the inception of the EISCAT KST, its successor, EISCAT_3D, is faced with a whole new situation. Solid state technology has advanced to the point where phased-array systems with thousands of individually driven elements are now the obvious choice for new radar systems and computational power many orders of magnitude larger than that of the EISCAT

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Acknowledgements

Ingemar Wolf and Lars-Göran Vanhainen, two colleagues from my time at the Kiruna site who both started with EISCAT before myself - Ingemar was actually working in the project from the very beginning in 1976 - have generously shared their own recollections of the early times. Mike Rietveld kindly provided photos of the capacitor bank and the crowbar.

Thanks to the referees for valuable corrections and comments.

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Appendix: About the author

1390 My background and route into EISCAT are quite different from most of my colleagues. In 1979, on
1395 completing my Ph.D. in nuclear physics at Uppsala University, I started to look for a job. At about
this time, EISCAT HQ was advertising in major Swedish newspapers. The EISCAT project looked
very technically attractive and I felt that maybe it could use me somehow; I had been trained as an
experimentalist, much of my work had been performed in an international environment at CERN,
our system contained a great deal of radio frequency technology - and I was also an enthusiastic
radio amateur who once dreamt of getting into radio astronomy professionally.

1400 That spring, I applied for a Scientific Programmer position at HQ, was interviewed and turned
down - but a few weeks later I got a phone call out of the blue from the Swedish Institute of Space
Physics. They had got hold of my application to EISCAT and liked it, I was invited to an interview
- and the next day I was offered a research engineer position in the satellite group! I jumped at the
opportunity and started on the job in November 1979, spending the first couple of months studying
up on plasma physics, but all the time I kept an eye out for any opportunities at EISCAT. When the
Kiruna Site Manager position was about to become vacant in 1981, I applied for it and got it,
1405 perhaps because I was the only applicant - Professor Hultqvist regarded my choice as a voluntary
demotion! The rest is history...

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